

Complex-Time (*Kime*)

At a given spatial location, x, complex time (kime) is defined by $\kappa = re^{i\varphi} \in \mathbb{C}$, where:

the magnitude represents the longitudinal events order (r > 0) and characterizes the longitudinal displacement in time, and

event phase ($-\pi \le \varphi < \pi$) is an angular displacement, or event direction

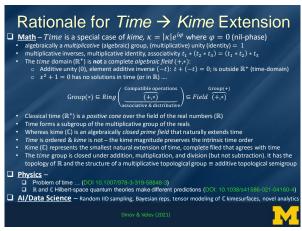
There are multiple alternative parametrizations of kime in the complex plane

space-kime manifold is $\mathbb{R}^3 \times \mathbb{C}$:

(x, k_1) and (x, k_2) have the same spacetime representation, but different spacekime coordinates,

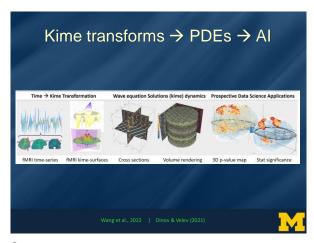
(x, k_1) and (x, k_2) have the same kime, but represent different spatial locations,

(x, k_2) and (x, k_3) have the same spatial-locations and kime-directions, but appear sequentially in order, $r_2 < r_1$.

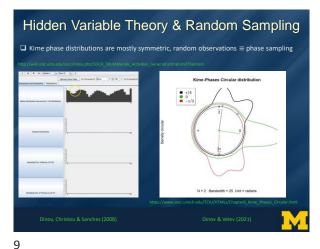


 $\begin{array}{c} \text{Ultrahyperbolic Wave Equation} - \\ \text{Cauchy Initial Data} \\ \hline $$ | & \text{Nonlocal constraints} \text{ yield the existence, uniqueness \& stability of local and global solutions to the ultrahyperbolic wave equation under Cauchy initial data ...} \\ \sum_{i=1}^{d_s} \partial_{x_i^2}^2 u \equiv \Delta_x u(x,\kappa) = \Delta_\kappa u(x,\kappa) \equiv \sum_{i=1}^{d_s} \partial_{k_i^2}^2 u, \\ \text{temporal Laplacian} & \text{temporal Laplacian} & \text{local constraints} \begin{bmatrix} u_o = u\left(\frac{x}{\kappa E p_s}, 0, \kappa_{-1}\right) = f(x,\kappa_{-1}) \\ u_i = \partial_{\kappa_i} u(x,0,\kappa_{-1}) = g(x,\kappa_{-1}) \\ \text{Initial conditions (Cauchy Data)} \\ \text{where } x = (x_1,x_2,\dots,x_d) \in \mathbb{R}^d \text{ and } \kappa = (x_1,x_2,\dots,x_d) \in \mathbb{R}^d \text{ are the Cartesian coordinates in the d_s space and d_i time dims.} \\ \text{Stable local solution over a Fourier frequency region defined by nonlocal constraints } |\xi| \geq |\eta_{-1}| : \\ \hat{u}\left(\xi,\frac{\kappa_1,\eta_{-1}}{\eta}\right) = \cos\left(2\pi\kappa_1\sqrt{|\xi|^2-|\eta_{-1}|^2}\right) \frac{\hat{u}_0(\xi,\eta_{-1})}{\hat{u}_0(\xi,\eta_{-1})} + \sin\left(2\pi\kappa_1\sqrt{|\xi|^2-|\eta_{-1}|^2}\right) \frac{\hat{u}_1(\xi,\eta_{-1})}{2\pi\sqrt{|\xi|^2-|\eta_{-1}|^2}}, \\ \text{where } \mathcal{F}\binom{u_o}{u_i} = \begin{pmatrix} \hat{u}_o(\xi,\eta_{-1}) \\ \hat{u}_1 \end{pmatrix} = \begin{pmatrix} \hat{u}_o(\xi,\eta_{-1}) \\ \hat{u}_1(\xi,\eta_{-1}) \end{pmatrix} = \begin{pmatrix} \hat{u}(\xi,\eta_{-1}) \\ \hat{u}_1(\xi,\eta_{-1}) \end{pmatrix} - \begin{pmatrix} \hat{u}(\xi,\eta_{-1}) \\ \hat{u}_1(\xi,\eta_{-1}) \end{pmatrix} + \frac{\hat{u}(\xi,\kappa_1,\eta_{-1})}{2\pi\sqrt{|\xi|^2-|\eta_{-1}|^2}}} \\ \hat{u}_1(\xi,\eta_{-1}) \end{pmatrix} = \mathcal{F}^{-1}(\hat{u})(x,\kappa) = \int_{D_i \times D_{k-1}} \hat{u}(\xi,\kappa_1,\eta_{-1}) \times e^{2\pi i(x,\xi)} \times e^{2\pi i(\kappa_{-1},\eta_{-1})} d\xi \, d\eta_{-1} \ . \\ \text{Wang et al., 2022} \qquad | \text{Dinov & Velev (2021)} \end{aligned}$





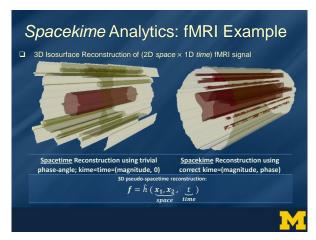
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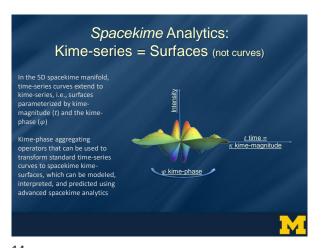
(Many) Spacekime Open Math Problems ☐ Ergodicity Let's look at particle velocities in the 4D Minkowski spacetime (X), a measure space where gas particles move spatially and evolve longitudinally in time. Let $\mu = \mu_X$ be a measure on X, $f(x,t) \in L^1(X,\mu)$ be an integrable function (e.g., <u>velocity</u> of a particle), and $\underline{T}: X \to X$ be a measure-preserving $\underline{\text{transformation}}$ at position $x \in \mathbb{R}^3$ and time $t \in \mathbb{R}^+$ A pointwise ergodic theorem argues that in a measure theoretic sense, the average of f (e.g., velocity) over all particles in the gas system at a fixed time, $ar f=E_t(f)=\int_{\mathbb{R}^3}f(x,t)d\mu_x$, will be equal to the average f of just one particle (x) over the entire time span $\tilde{f} = \lim_{n \to \infty} \left(\frac{1}{n} \sum_{m=0}^{n} f(T^m x) \right)$, i.e., (show) $\bar{f} \equiv \tilde{f}$. The spatial probability measure is denoted by μ_x and the transformation $T^m x$ represents the dynamics (time evolution) of the particle starting with an initial spatial location $T^o x = x$ <u>Investigate</u> the ergodic properties of various transformations in the 5D <u>spacekime</u>: $\bar{f} \equiv E_{\kappa}(f) = \frac{1}{\mu_{x}(X)} \int f\left(x, \underline{t}, \underline{\phi}\right) d\mu_{x} \cong$ space averaging

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Physics	Data/Neuro Sciences
A <u>particle</u> is a small localized object that permits observations and characterization of its physical or chemical properties An <u>observable</u> a dynamic variable about	An <u>object</u> is something that exists by itself, actually or potentially, concretely or abstractly, physically or incorporeal (e.g., person, subject, etc.) A <u>feature</u> is a dynamic variable or an attribute about a
particles that can be measured Particle state is an observable particle characteristic (e.g., position, momentum)	object that can be measured Datum is an observed quantitative or qualitative value an instantiation, of a feature
Particle system is a collection of independent particles and observable characteristics, in a closed system	Problem, aka Data System, is a collection of independent objects and features, without necessarily being associated with a priori hypotheses
Wave-function	Inference-function
Reference-Frame transforms (e.g., Lorentz)	Data transformations (e.g., wrangling, log-transform)
State of a system is an observed	Dataset (data) is an observed instance of a set of
measurement of all particles – wavefunction A particle system is computable if (1) the entire system is logical, consistent, complete and (2) the unknown internal states of the system don't influence the computation (wavefunction, intervals, probabilities, etc.)	datum elements about the problem system, $O = (X, Y)$ <u>Computable data object</u> is a very special representation of a dataset which allows direct application of computational processing, modeling, analytics, or inference based on the observed dataset



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Spacetime Time-series ⇒ Spacekime Kimesurfaces ⇒ TLM

Difference for ON & OFF Kime-Surface/Kime-Series at a fixed voxel location

Difference for ON & OFF Kime-Surface/Kime-Series at a fixed voxel location

Trace 7

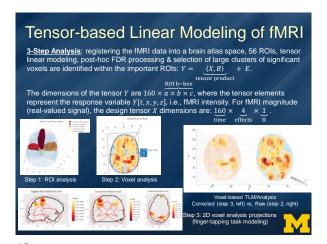
forecasted time series

95% upper bound
95% lower bound
ordinal time series

Time

Zhang et al., 2022 | Dinov & Velev (2021)

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Bayesian Inference Representation

Suppose we have a single spacetime observation $X = \{x_{l_0}\} \sim p(x \mid \gamma)$ and $\gamma \sim p(\gamma \mid \varphi = \text{phase})$ is a process parameter (or vector) that we are trying to estimate.

Spacekime analytics aims to make appropriate inference about the process X.

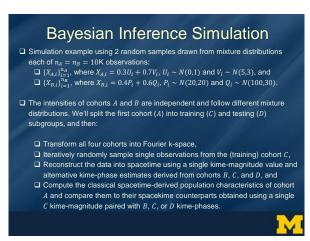
The <u>sampling distribution</u>, $p(x \mid \gamma)$, is the distribution of the observed data X conditional on the parameter γ and the <u>prior distribution</u>, $p(\gamma \mid \varphi)$, of the parameter γ before the data X is observed, $\varphi = \text{phase aggregator}$.

Assume that the hyperparameter (vector) φ , which represents the kime-phase estimates for the process, can be estimated by $\hat{\varphi} = \varphi'$.

Such estimates may be obtained from an oracle (model distribution), approximated using similar datasets, acquired as phases from samples of analogous processes, derived via some phase-aggregation strategy, or computed via Laplace transform.

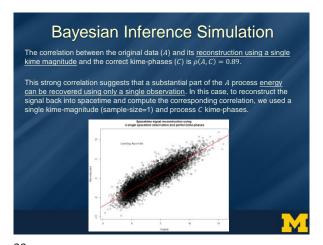
Let the <u>posterior distribution</u> of the parameter γ given the observed data $X = \{x_{l_0}\}$ be $p(\gamma|X,\varphi')$ and the process parameter distribution of the kime-phase hyperparameter vector φ be $\gamma \sim p(\gamma \mid \varphi)$.

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Bayesian Inference Simulation Summary statistics for the original process (cohort A) and the corresponding values of their counterparts computed using the spacekime reconstructed signals based on kime-phases of cohorts *B*, *C*, and *D*. The estimates for the latter three cohorts correspond to reconstructions using a single spacetime observation (i.e., single kime-magnitude) and alternative kime-phase priors (in this case, kime-phases derived from cohorts B, C, and D). Spacetime | Spacekime Reconstructions (single kime-magnitude) (A) (B) (C) (D) hase=Tru -2.98116 e=Indeper -2.69808 -0.76453 Original -2.38798 e=Diff. Process -3.798440 -0.89359 -0.636799 -0.76765 0.03311 0.009279 0.00000 0.00000 0.00000 3rd Qu 0.645119 0.72795 3.61346 3.986702 3.64800 3.22987 0.001021943 0.2372526 -0.68176 0.2149918 B A

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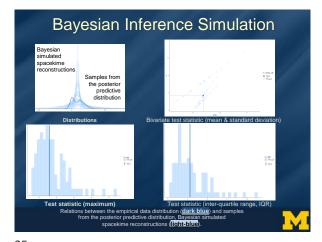
Bayesian Inference Simulation

Let's demonstrate the Bayesian inference corresponding to this spacekime data analytic problem using a simulated bimodal experiment: $X_A = 0.3U + 0.7V, \text{ where } U \sim N(0,1) \text{ and } V \sim N(5,3)$ Specifically, we will illustrate the Bayesian inference using repeated single spacetime observations from cohort $A, X = \{x_{i_0}\}$, and varying kime-phase priors $(\theta = \text{phase aggregator})$ obtained from cohorts B, C, or D, using different posterior predictive distributions.

Relations between the empirical data distribution (dark blue) and samples from the posterior predictive distribution, representing Bayesian simulated spacekime reconstructions (light-blue). The derived Bayesian estimates do not perfectly match the empirical distribution of the simulated data, yet there is clearly information encoding that is captured by the spacekime data reconstructions.

This signal compression can be exploited by subsequent model-based or model-free data analytic strategies for retrospective prediction, prospective forecasting, ML classification, Al derived clustering, and other spacekime inference methods.

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Spacekime Analytics: Demos

Tutorials
https://TCIU.predictive.space
https://SpaceKime.org

R Package
https://cran.rstudio.com/web/packages/TCIU

GitHub
https://github.com/SOCR/TCIU

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