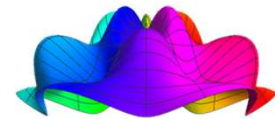


# Complex-time Representation of Repeated Measurement Longitudinal Data & Space-kime Analytics

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## Outline

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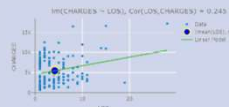
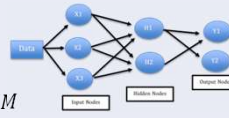
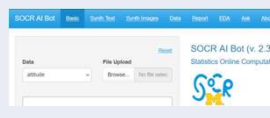
- Data Science  $\equiv$  need-based information compression & expansion
- Complex-time (*kime*) & rationale
- Kime-phase, random sampling & Heisenberg's Uncertainty
- Solutions of ultrahyperbolic wave equations
- Open spacekime problems
- Data science applications
- Bayesian formulation of spacekime inference
- Resources, live demo links & prospective DS R&D, education & practice
- Part 2: Hands-on Spacekime Analytics Tutorial (Demos)*



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## Duality of Evidence-based Scientific Discovery

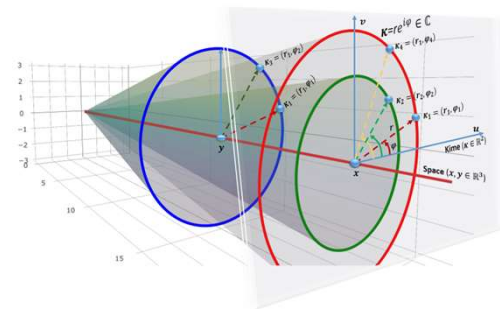
experimental → theoretical → computational → data sciences

Mapping Examples	Analysis Observables/Data → Compact Models	Synthesis Compact Models → (simulated, actionable info)
1. Lossless Math Transforms	(A.1.1) <u>Linear transform</u> , $L: V \rightarrow W$ , e.g., 2D rigid body $L = \begin{pmatrix} \cos \theta & -\sin \theta \\ \sin \theta & \cos \theta \end{pmatrix}: \mathbb{R}^2 \xrightarrow{\text{rotation}} \mathbb{R}^2$ (A.1.2) <u>Fourier transform</u> : $\hat{f}(\omega) = \int_{-\infty}^{\infty} f(x)e^{-i2\pi\omega x} dx$	(S.1.1) <u>Inverse linear transform</u> , $L^{-1}: W \rightarrow V$ , e.g., $L^{-1} = \begin{pmatrix} \cos \theta & \sin \theta \\ -\sin \theta & \cos \theta \end{pmatrix}: \mathbb{R}^2 \xrightarrow{\text{rotation}} \mathbb{R}^2, \quad LL^{-1} \equiv \mathbb{I}$ (S.1.2) <u>Inverse Fourier (IFT)</u> : $f(x) = \int_{-\infty}^{\infty} \hat{f}(\omega)e^{i2\pi\omega x} d\omega$
2. DNA	(A.2.1) <u>DNA Packing</u> in Chromatin Fiber Chromosomes contain enormously long linear DNA molecules associated with proteins that fold and pack the fine DNA double helix into a <i>tight compact structure</i>	(S.2.1) <u>DNA Unpacking</u> The process of unfolding the DNA from the chromosome to support the processes of <u>gene expression</u> , <u>DNA replication</u> , and <u>DNA repair</u>
3. Lossy Data/Stats Science	(A.3.1) <u>Info Compression</u> , e.g., linear models $Y = 4582.70 + 212.29 X$ Data $\xrightarrow{\text{assumptions}}$ Model 	(S.3.1) <u>Information Inflation, Simulation &amp; Generation</u> , e.g., forecasting, regression, interpolation, extrapolation (predict & classify new data): Input $\xrightarrow{\text{mod}}$ Output
4. Artificial & Augmented Intelligence	(A.4.1) <u>Building, Fitting &amp; Training</u> large foundational, generative & deep network AI models Data $\xrightarrow{\text{human+infrast}}$ GAIM 	(S.4.1) Generative Artificial Intelligence Modeling (GAIM) Human Prompt $\xrightarrow{\text{GAIM}}$ Result 

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## Complex-Time (Kime)

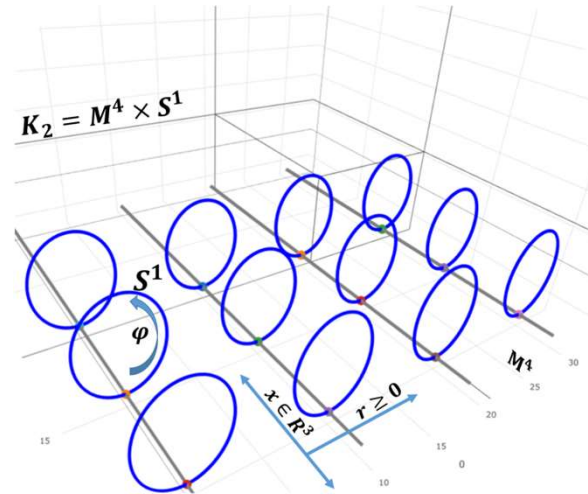
- ❑ At a given spatial location,  $x$ , complex time (*kime*) is defined by  $\kappa = re^{i\varphi} \in \mathbb{C}$ , where:
  - ❑ the magnitude represents the longitudinal events order ( $r > 0$ ) and characterizes the longitudinal displacement in time, and
  - ❑ event phase ( $-\pi \leq \varphi < \pi$ ) is an angular displacement, event direction, or random sampling index
- ❑ There are multiple alternative parametrizations of kime in the complex plane
- ❑ Space-kime manifold is  $\mathbb{R}^3 \times \mathbb{C}$ :
  - ❑  $(x, k_1)$  and  $(x, k_4)$  have the same spacetime representation, but different spacekime coordinates,
  - ❑  $(x, k_1)$  and  $(y, k_1)$  share the same kime, but represent different spatial locations,
  - ❑  $(x, k_2)$  and  $(x, k_3)$  have the same spatial-locations and kime-directions, but appear sequentially in order,  $r_2 < r_1$ .



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## Historical Background: Kaluza-Klein Theory

- Theodor Kaluza (1921) developed a math extension of the classical general relativity theory to 5D. This included the metric, the field equations, the equations of motion, the stress-energy tensor, and the cylinder condition. Physicist Oskar Klein (1926) interpreted Kaluza's 3D+2D theory in quantum mechanical space and proposed that the fifth dimension was curled up and microscopic.
- The topology of the 5D Kaluza-Klein spacetime is  $K_2 \cong M^4 \times S^1$ , where  $M^4$  is a 4D Minkowski spacetime and  $S^1$  is a circle (non-traversable).



**M**

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## AI & Spacetime Analytics

### Rationale for *Time* $\Rightarrow$ *Kime* Extension

**Math** – *Time* is a special case of *kime*,  $\kappa = |\kappa|e^{i\varphi}$  where  $\varphi = 0$

**Time** ( $\mathbb{R}^+$ ) is a subgroup of the multiplicative Reals group

Whereas **kime** ( $\mathbb{C}$ ) is an algebraically closed prime field that naturally extends time

*Time* is ordered but *kime* is not!

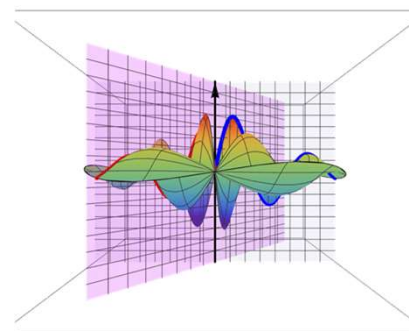
*Kime* ( $\mathbb{C}$ ) represents the smallest natural extension of time, as a complete field that agrees with time

**Physics** –

Problem of time ... (DOI 10.1007/978-3-319-58848-3)

$\mathbb{R}$  and  $\mathbb{C}$  Hilbert-space quantum theories make different predictions (DOI: 10.1038/s41586-021-04160-4)

**Bio AI/Data Science** – Random IID sampling, Bayesian reps, tensor modeling of  $\mathbb{C}$  kimesurfaces, novel analytics



Wesson (2004, 2010)  
Dinov & Velez (2021)  
Wang et al. (2022)  
Zhang et al. (2023)  
Dinov & Shen (2024)

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## Uncertainty in 5D Spacekime

In 5D space-time,  $\Omega = \left(\frac{l}{L}\right)^2$  is the conformal factor, and  $L$  is a constant length defined in terms of the cosmological constant

$$\Lambda = -\epsilon \frac{3}{L^2}. \text{ In the metric signature } (+, -, -, -),$$

$\Lambda > 0$  for a spacelike extra coordinate, and  $\Lambda < 0$  for a time-like extra 5<sup>th</sup> coordinate,  $x^\mu$  is the  $(D-1)$  spacetime location, and  $l$  is the extra kime dimension.

The *canonical spacekime metric* is:

$$dS^2 = \frac{l^2}{L^2} \sum_0^{D-2} \sum_0^{D-2} g_{\alpha\beta}(x^\mu, l) dx^\alpha dx^\beta + \epsilon dl^2 \quad (5D \text{ Spacekime line element and metric})$$

The 4D components of the spacekime equations of motion can be written explicitly in terms of the fifth force  $f^\mu$  measured in units of inertia mass, i.e., assuming  $m = 1$ :

$$\frac{du^\mu}{ds} + \sum_0^3 \sum_0^3 \Gamma_{\beta\gamma}^\mu u^\beta u^\gamma = f^\mu, \quad f^\mu \equiv \sum_{\alpha=0}^3 \sum_{\beta=0}^3 \left( -g^{\mu\alpha} + \frac{1}{2} u^\mu u^\alpha \right) \frac{dl}{ds} \frac{dx^\beta}{ds} \frac{\partial g_{\alpha\beta}}{\partial l}$$

The 5D component of the spacekime equation of motion is:

$$\frac{d^2 l}{ds^2} - \frac{2}{l} \left( \frac{dl}{ds} \right)^2 - \frac{l}{L^2} = \frac{1}{2} \left[ \frac{l^2}{L^2} + \left( \frac{dl}{ds} \right)^2 \right] \sum_{\alpha=0}^3 \sum_{\beta=0}^3 u^\alpha u^\beta \frac{\partial g_{\alpha\beta}}{\partial l} \quad \& \quad f_\parallel^\mu = -\frac{1}{2} u^\mu \sum_{\alpha=0}^3 \sum_{\beta=0}^3 \left( \frac{\partial g_{\alpha\beta}}{\partial l} u^\alpha u^\beta \right) \frac{dl}{ds}$$

In 5D spacekime, geodesic motion is perturbed by an extra **5<sup>th</sup> force**  $f^\mu = f_\perp^\mu + f_\parallel^\mu$ , where

- $f_\perp^\mu$  is normal to the 4-velocity  $u_\mu$ , similar to other conventional forces, and  $f_\perp^\mu u_\mu = 0$
- $f_\parallel^\mu$  is parallel to the 4-velocity  $u_\mu$ , has no analog in 4D spacetime, and  $f_\parallel^\mu u_\mu \neq 0$



Wesson (2004, 2010) | Wesson & Overduin (2018) | Dinov & Velev (2021)

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## Uncertainty in 5D Spacekime

- Assuming  $m = 1, c = 1$ , near the foliation leaf membrane hypersurface, we have

$$\langle dp | dx \rangle = \sum_{\mu=0}^3 dp^\mu dx_\mu = L \left( \frac{dl}{l-l_0} \right)^2 = \frac{h}{m c} \left( \frac{dl}{l-l_0} \right)^2 \sim h$$

derived from 5D Einstein deterministic field equ's  $\Rightarrow$  uncertainty principle in 4D Minkowski spacetime

- In spacetime, Heisenberg's uncertainty is due to lack of sufficient information about the 2<sup>nd</sup> kime dimension,  $l$ .
- In Minkowski 4D spacetime, the lack of kime-phase information naturally leaves one degree of freedom (**DoF**) in the system, which appears as Heisenberg's uncertainty.
- **In Bioinfo/Biostatistics, Data Science, ML/AI & longitudinal analysis, this extra DoF represents process stochasticity – random sampling from an underlying probability distribution**
- Spacekime formulation of the 4D spacetime observation of the Heisenberg's principle also supports the de Broglie-Bohm theory, which provides an explicit deterministic model of a system configuration and its corresponding wavefunction
- 4D probabilistic spacetime is a spacekime embedding with an added degrees of freedom
- Bell's theorem suggests that any deterministic hidden-variable theory, which is consistent with quantum mechanics predictions, has to be non-local. This implies the existence of instantaneous, faster than the speed of light, interactions between particles that are significantly separated in 3D space (non-local relations).



Wesson (2004, 2010) | Bell (1964) | Dinov & Velev (2021)

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## Ultrahyperbolic Wave Equation – Cauchy Initial Data

- **Nonlocal constraints** yield the existence, uniqueness & stability of local and global solutions to the ultrahyperbolic wave equation under Cauchy initial data ...

$$\underbrace{\sum_{i=1}^{d_s} \partial_{x_i}^2 u \equiv \Delta_x u(\mathbf{x}, \boldsymbol{\kappa})}_{\text{spatial Laplacian}} = \underbrace{\Delta_{\boldsymbol{\kappa}} u(\mathbf{x}, \boldsymbol{\kappa}) \equiv \sum_{i=1}^{d_t} \partial_{\kappa_i}^2 u}_{\text{temporal Laplacian}}, \quad \begin{cases} u_0 = u(\mathbf{x}, 0, \boldsymbol{\kappa}_{-1}) = f(\mathbf{x}, \boldsymbol{\kappa}_{-1}) \\ u_1 = \partial_{\kappa_1} u(\mathbf{x}, 0, \boldsymbol{\kappa}_{-1}) = g(\mathbf{x}, \boldsymbol{\kappa}_{-1}) \end{cases}$$

initial conditions (Cauchy Data)

where  $\mathbf{x} = (x_1, x_2, \dots, x_{d_s}) \in \mathbb{R}^{d_s}$  and  $\boldsymbol{\kappa} = (\kappa_1, \kappa_2, \dots, \kappa_{d_t}) \in \mathbb{R}^{d_t}$  are the Cartesian coordinates in the  $d_s$  space and  $d_t$  time dims.

Stable local solution over a Fourier frequency region defined by **nonlocal constraints**  $|\boldsymbol{\xi}| \geq |\boldsymbol{\eta}_{-1}|$  :

$$\hat{u}(\boldsymbol{\xi}, \kappa_1, \boldsymbol{\eta}_{-1}) = \cos\left(2\pi \kappa_1 \sqrt{|\boldsymbol{\xi}|^2 - |\boldsymbol{\eta}_{-1}|^2}\right) \underbrace{\hat{u}_0(\boldsymbol{\xi}, \boldsymbol{\eta}_{-1})}_{c_1} + \sin\left(2\pi \kappa_1 \sqrt{|\boldsymbol{\xi}|^2 - |\boldsymbol{\eta}_{-1}|^2}\right) \underbrace{\frac{\hat{u}_1(\boldsymbol{\xi}, \boldsymbol{\eta}_{-1})}{2\pi \sqrt{|\boldsymbol{\xi}|^2 - |\boldsymbol{\eta}_{-1}|^2}}}_{c_2},$$

$$\text{where } \mathcal{F} \begin{pmatrix} u_0 \\ u_1 \end{pmatrix} = \begin{pmatrix} \hat{u}_0 \\ \hat{u}_1 \end{pmatrix} = \begin{pmatrix} \hat{u}_0(\boldsymbol{\xi}, \boldsymbol{\eta}_{-1}) \\ \hat{u}_1(\boldsymbol{\xi}, \boldsymbol{\eta}_{-1}) \end{pmatrix} = \begin{pmatrix} \hat{u}(\boldsymbol{\xi}, \boldsymbol{\eta}_{-1}) \\ \partial_{\kappa_1} \hat{u}(\boldsymbol{\xi}, \boldsymbol{\eta}_{-1}) \end{pmatrix}.$$

$$u(\mathbf{x}, \kappa_1, \boldsymbol{\kappa}_{-1}) = \mathcal{F}^{-1}(\hat{u})(\mathbf{x}, \boldsymbol{\kappa}) = \int_{\tilde{D}_s \times \tilde{D}_{t-1}} \hat{u}(\boldsymbol{\xi}, \kappa_1, \boldsymbol{\eta}_{-1}) \times e^{2\pi i \langle \mathbf{x}, \boldsymbol{\xi} \rangle} \times e^{2\pi i \langle \kappa_{-1}, \boldsymbol{\eta}_{-1} \rangle} d\boldsymbol{\xi} d\boldsymbol{\eta}_{-1}.$$



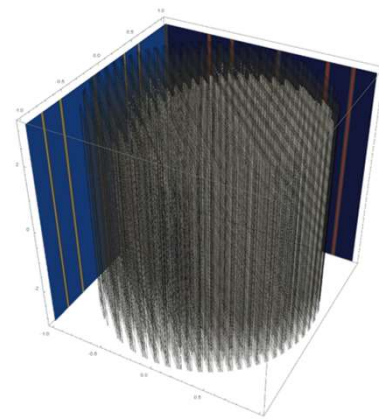
Craig & Weinstein (2008) | Wang et al. (2022) | Dinov & Velev (2021)

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## Ultrahyperbolic Wave Equation – Cauchy Initial Data

- **Math Generalizations:**

Derived **other spacekime concepts**: law of addition of velocities, energy-momentum conservation law, stability conditions for particles moving in spacekime, conditions for nonzero rest particle mass, causal structure of spacekime, and **solutions** of the ultrahyperbolic wave equation under Cauchy initial data ...



Wang et al., 2022 | Dinov & Velev (2021)

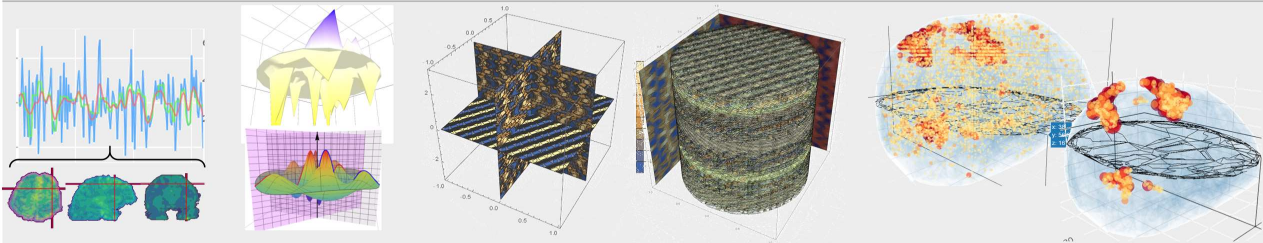
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# Longitudinal Bio Data $\Rightarrow$ Kime-Transforms $\Rightarrow$ PDEs $\Rightarrow$ AI

Time  $\rightarrow$  Kime Transformation

Wave equation Solutions (kime) dynamics

Prospective Data Science Applications



fMRI time-series

fMRI kime-surfaces

Cross sections

Volume rendering

3D p-value map

Stat significance

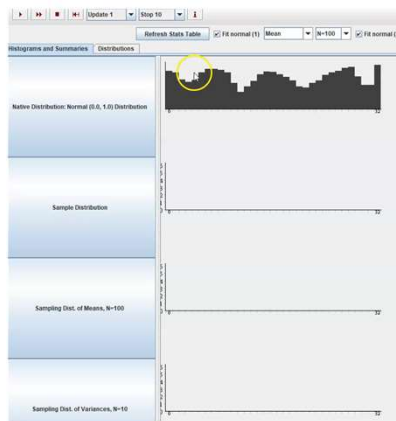


Wang et al., 2022 | Dinov & Velev (2021)

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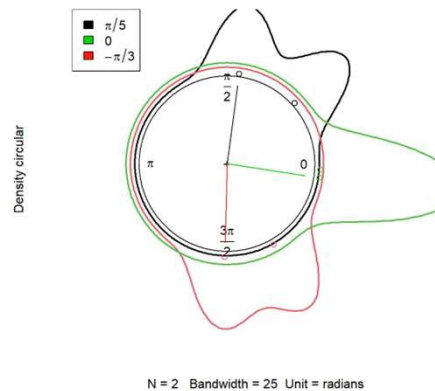
## Random Sampling & Kime-Phase Paradigm

Kime phase distributions are mostly symmetric, random observations  $\equiv$  phase sampling



[https://wiki.socr.umich.edu/index.php/SOCR\\_EduMaterials\\_Activities\\_GeneralCentralLimitTheorem](https://wiki.socr.umich.edu/index.php/SOCR_EduMaterials_Activities_GeneralCentralLimitTheorem)

Kime-Phases Circular distribution



[https://www.socr.umich.edu/TCIU/HTMLs/Chapter6\\_Kime\\_Phases\\_Circular.html](https://www.socr.umich.edu/TCIU/HTMLs/Chapter6_Kime_Phases_Circular.html)



Dinov, Christou & Sanchez (2008)

Dinov & Velev (2021)

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# (Many) Spacetime Open Math Problems

## Ergodicity

Let's look at particle velocities in the 4D Minkowski spacetime ( $X$ ), a measure space where gas particles move spatially and evolve longitudinally in time. Let  $\mu = \mu_x$  be a measure on  $X$ ,  $f(x, t) \in L^1(X, \mu)$  be an integrable function (e.g., velocity of a particle), and  $T: X \rightarrow X$  be a measure-preserving transformation at position  $x \in \mathbb{R}^3$  and time  $t \in \mathbb{R}^+$ .

A pointwise ergodic theorem argues that in a measure theoretic sense, the average of  $f$  (e.g., velocity) over all particles in the gas system at a fixed time,  $\bar{f} = \mathbb{E}_t(f) = \int_{\mathbb{R}^3} f(x, t) d\mu_x$ , will be equal to the average  $f$  of just one particle ( $x$ ) over the entire time span,

$$\bar{f} \equiv \mathbb{E}_x(f) = \lim_{n \rightarrow \infty} \left( \frac{1}{n} \sum_{m=0}^{n-1} f(T^m x) \right), \text{ i.e., (show) } \bar{f} \equiv \bar{f}.$$

The spatial probability measure is denoted by  $\mu_x$  and the transformation  $T^m x$  represents the dynamics (time evolution) of the particle starting with an initial spatial location  $T^0 x = x$ .

Investigate the ergodic properties of various transformations in the 5D **spacetime**:

$$\underbrace{\bar{f} \equiv \mathbb{E}_x(f) = \frac{1}{\mu_x(X)} \int f(x, t, \phi) d\mu_x}_{\text{space averaging}} \stackrel{?}{=} \underbrace{\lim_{t \rightarrow \infty} \left( \frac{1}{t} \sum_{m=0}^{t-1} \left( \int_{-\pi}^{+\pi} f(T^m x, t, \phi) d\Phi \right) \right)}_{\text{kime averaging}} = \mathbb{E}_x(f) \equiv \bar{f}$$



Dinov & Velev (2021)

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# Mathematical-Physics $\Rightarrow$ Bio-Data Science & AI

Physics	Bio-Data Sciences
A <b>particle</b> is a small localized object that permits observations and characterization of its physical or chemical properties	An <b>object</b> is something that exists by itself, actually or potentially, concretely or abstractly, physically or incorporeal (e.g., person, subject, etc.)
An <b>observable</b> a dynamic variable about particles that can be measured	A <b>feature</b> is a dynamic variable or an attribute about an object that can be measured
Particle <b>state</b> is an observable particle characteristic (e.g., position, momentum)	<b>Datum</b> is an observed quantitative or qualitative value, an instantiation, of a feature
Particle <b>system</b> is a collection of independent particles and observable characteristics, in a closed system	<b>Problem</b> , aka Data System, is a collection of independent objects and features, without necessarily being associated with a priori hypotheses
<b>Wave-function</b>	<b>Inference-function</b>
Reference-Frame <b>transforms</b> (e.g., Lorentz)	Data <b>transformations</b> (e.g., wrangling, log-transform)
<b>State of a system</b> is an observed measurement of all particles ~ wavefunction	<b>Dataset (data)</b> is an observed instance of a set of datum elements about the problem system, $\mathcal{O} = \{X, Y\}$
A <b>particle system is computable</b> if (1) the entire system is logical, consistent, complete and (2) the unknown internal states of the system don't influence the computation (wavefunction, intervals, probabilities, etc.)	<b>Computable data object</b> is a very special representation of a dataset which allows direct application of computational processing, modeling, analytics, or inference based on the observed dataset
⋮	⋮



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# Mathematical-Physics $\Rightarrow$ Bio-Data Science & AI

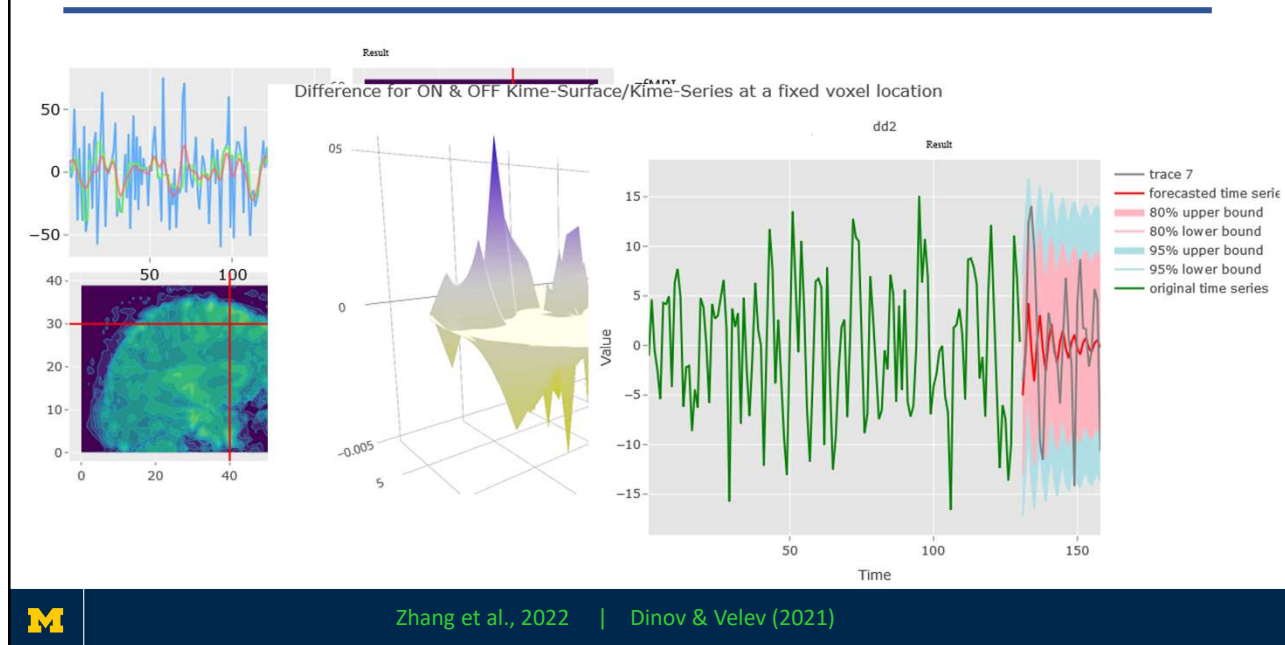
Physics	Data Science
<p><u>Wavefunction</u></p> <p>Wave equ problem:</p> $\left(\frac{\partial^2}{\partial x^2} - \frac{1}{v^2} \frac{\partial^2}{\partial t^2}\right)\psi(x, t) = 0$ <p>Complex Solution:  <math>\psi(x, t) = Ae^{i(kx - \omega t)}</math>                      represents a traveling wave,                      where <math>\left \frac{\omega}{k}\right  = v</math>.</p>	<p><u>Inference function</u> - describing a solution to a specific data analytic system (a problem). Examples:</p> <ul style="list-style-type: none"> <li>A <u>linear (GLM) model</u> represents a solution of a prediction inference problem, <math>Y = X\beta</math>, where the inference function quantifies the effects of all independent features (<math>X</math>) on the dependent outcome (<math>Y</math>), data: <math>O = \{X, Y\}</math>:  <math display="block">\psi(O) = \psi(X, Y) \Rightarrow \hat{\beta} = \hat{\beta}^{OLS} = (X X)^{-1} \langle X Y \rangle = (X^T X)^{-1} X^T Y.</math></li> <li>A non-parametric, <u>non-linear</u>, alternative inference is SVM classification. If <math>\psi_x \in H</math>, is the lifting function <math>\psi: R^n \rightarrow R^d</math> (<math>\psi: x \in R^n \rightarrow \tilde{x} = \psi_x \in H</math>), where <math>\eta \ll d</math>, the kernel <math>\psi_x(y) = \langle x y \rangle: O \times O \rightarrow R</math> transforms non-linear to linear separation, the observed data <math>O_i = \{x_i, y_i\} \in R^n</math> are lifted to <math>\psi_{O_i} \in H</math>. The SVM prediction operator is the weighted sum of the kernel functions at <math>\psi_{O_i}</math>, where <math>\beta^*</math> is a solution to the SVM regularized optimization:  <math display="block">\langle \psi_{O_i}   \beta^* \rangle_H = w^T x + b = \sum_{i=1}^n p_i^* \langle \psi_{O_i}   \psi_{O_i} \rangle_H + b,</math> <math display="block">\min_{w \in R^d, \xi \in R^+} \left( \underbrace{\ w\ ^2}_{\text{regularizer}} + C \underbrace{\sum_{i=1}^m \xi_i}_{\text{fidelity}} \right), y^{(i)}(w^T x^{(i)} + b) \geq 1 - \xi_i, \xi_i \geq 0</math>                     The dual weight coefficients, <math>p_i^*</math>, are multiplied by the label corresponding to each training instance, <math>\{y^{(i)}\}</math>.                      Inference always depends on the (input) data; however, it does not have 1-1 and onto bijective correspondence with the data, since the inference function quantifies predictions probabilistically.</li> </ul>



GLM/SVM: [https://DSPA2\\_predictive.space](https://DSPA2_predictive.space) | Dinov, Springer (2018, 2023)

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# Spacetime Time-series $\Rightarrow$ Spacekime Kimesurfaces $\Rightarrow$ TLM

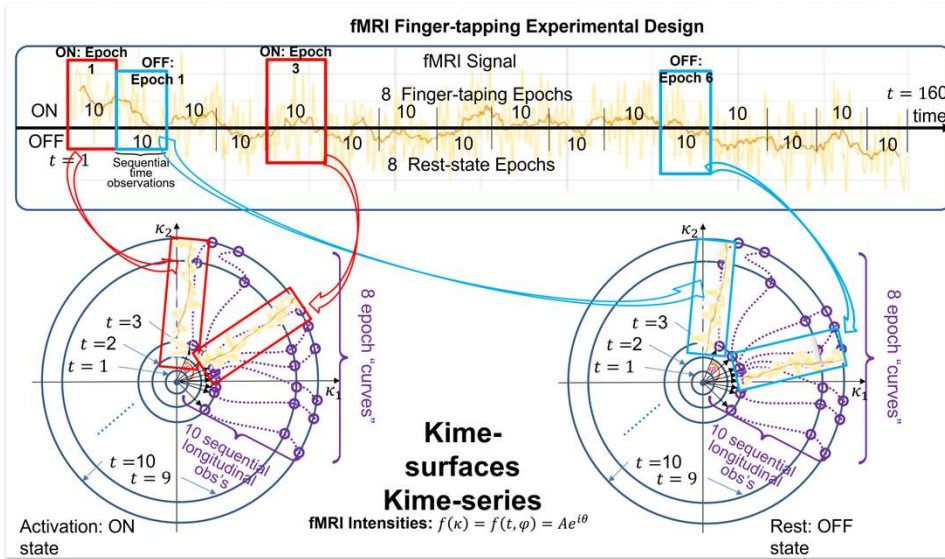


Zhang et al., 2022 | Dinov & Velev (2021)

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# Mapping Longitudinal Bio-Data (Time-series) $\Rightarrow$ Kime-Surfaces



Zhang et al., 2022 | Dinov & Velev (2021)

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# (Analytic) Mapping Bio Time-series $\Rightarrow$ Kime-surfaces

Apply the ILT ( $\mathcal{L}^{-1}$ ) to reconstruct a time-series,  $f(t) = \mathcal{L}^{-1}(F)(t)$ :

$$F(z) = \mathcal{L}(f) = \frac{1}{z+1} + \frac{1}{z^2+1} \times \frac{z}{z^2+1} + \frac{1}{z^2}$$

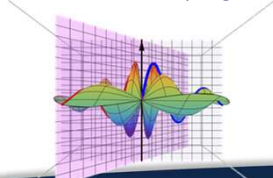
$F_1(z) = \mathcal{L}(f_1(t) = e^{-t})$     $F_2(z) = \mathcal{L}(f_2(t) = \sin(t))$     $F_3(z) = \mathcal{L}(f_3(t) = \cos(t))$     $F_4(z) = \mathcal{L}(f_4(t) = t)$

$$f(t) = \mathcal{L}^{-1}(F) = \mathcal{L}^{-1}(F_1 + F_2 * F_3 + F_4) = \mathcal{L}^{-1}(F_1) + \left( \frac{\mathcal{L}^{-1}(F_2) * \mathcal{L}^{-1}(F_3)}{\text{convolution}} \right) + \mathcal{L}^{-1}(F_4) =$$

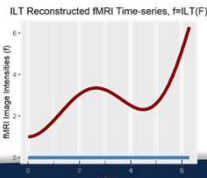
$$\mathcal{L}^{-1}(\mathcal{L}(f_1))(t) + \left( \mathcal{L}^{-1}(\mathcal{L}(f_2)) * \mathcal{L}^{-1}(\mathcal{L}(f_3)) \right) (t) + \mathcal{L}^{-1}(\mathcal{L}(f_4))(t),$$

$$f(t) = \mathcal{L}^{-1}(F)(t) = f_1(t) + (f_2 * f_3)(t) + f_4(t) = e^{-t} + \int_0^t \sin(\tau) \times \cos(t - \tau) d\tau + t = t + e^{-t} + \frac{t \sin(t)}{2}.$$

Repeated Longitudinal Data Sampling

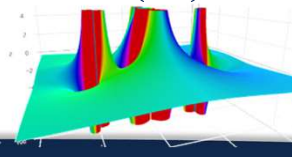


$$f(t) = \mathcal{L}^{-1}(\mathcal{L}(f))(t)$$



Kime-Surface, Height=Re(F), Color=Im(F)

$$F = \mathcal{L}(f) = 1/(z+1) + (1/(z^2+1)) * (z/(z^2+1)) + 1/(z^2)$$

$$F(z) = \mathcal{L}(f(\cdot))(z)$$


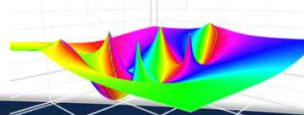
Inverse stereographic projection

Regularized Kime-Surface (LT)

Height=Zenith(F), Color=Azimuth(F)

$$F = \mathcal{L}(f) = 1/(z+1) + (1/(z^2+1)) * (z/(z^2+1)) + 1/(z^2)$$

$$\text{Reg}(F)(z) = \text{Reg}(\mathcal{L}(f))(z)$$



Shen et al., 2024 | Zhang et al., 2022 | Dinov & Velev (2021)

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## Example: Tensor-based Linear Modeling of fMRI

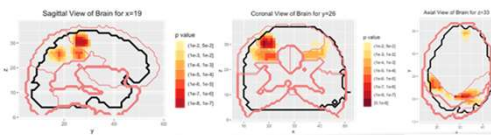
**3-Step Analysis:** registering the fMRI data into a brain atlas space, 56 ROIs, tensor linear modeling, post-hoc FDR processing & selection of large clusters of significant voxels are identified within the important ROIs:

$$Y = \underbrace{\langle X, B \rangle}_{\text{tensor product}} + E$$

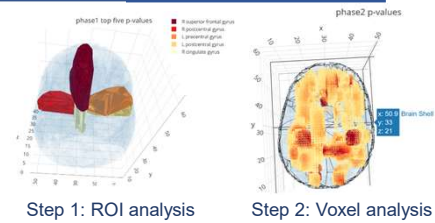
time      ROI b-bo

The dimensions of the time-tensor  $Y$  are  $160 \times a \times b \times c$ , where the tensor elements represent the response variable  $Y[t, x, y, z]$ , i.e., fMRI intensity. For fMRI magnitude (real-valued signal), the design kime-tensor  $X$  dimensions are:

$$\underbrace{10 * 8}_{\text{Kime}(\text{Time} * e^{i * \text{Repeat}})} \times \underbrace{\text{State}}_{\text{Stim vs. Rest (2)}} \times \underbrace{4}_{\text{effects}} \times \underbrace{1}_{\mathbb{R}}$$

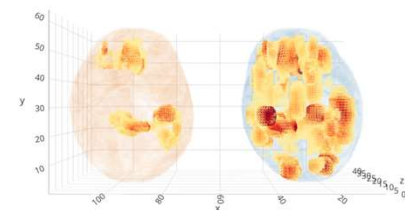


Step 3: 2D voxel analysis projections (finger-tapping task modeling)



Step 1: ROI analysis

Step 2: Voxel analysis



Voxel-based TLM/Analysis  
Corrected (step 3, left) vs. Raw (step 2, right)



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## Bayesian Inference Representation

- ❑ Suppose we have a single spacetime observation  $X = \{x_{i_o}\} \sim p(x | \gamma)$  and  $\gamma \sim p(\gamma | \varphi = \text{phase})$  is a process parameter (or vector) that we are trying to estimate.
- ❑ Spacekime analytics aims to make appropriate inference about the process  $X$ .
- ❑ The sampling distribution,  $p(x | \gamma)$ , is the distribution of the observed data  $X$  conditional on the parameter  $\gamma$  and the prior distribution of the parameter  $\gamma$  before the observing the data is  $p(\gamma | \varphi)$ , where  $\varphi = \text{phase aggregator}$ .
- ❑ Assume that the hyperparameter (vector)  $\varphi$ , which represents the kime-phase estimates for the process, can be estimated by  $\hat{\varphi} = \varphi'$ .
- ❑ Such estimates may be obtained from an oracle (model distribution), approximated using similar datasets, acquired as phases from samples of analogous processes, derived via some phase-aggregation strategy, or analytically computed (e.g., via Laplace transform).
- ❑ Let the posterior distribution of the parameter  $\gamma$  given the observed data  $X = \{x_{i_o}\}$  be  $p(\gamma | X, \varphi')$  and the process parameter distribution of the kime-phase hyperparameter vector  $\varphi$  be  $\gamma \sim p(\gamma | \varphi)$ .



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## Bayesian Inference Representation

- We can formulate spacetime inference as a Bayesian parameter estimation problem:

$$\underbrace{p(\gamma|X, \varphi')}_{\text{posterior distribution}} = \frac{p(\gamma, X, \varphi')}{p(X, \varphi')} = \frac{p(X|\gamma, \varphi') \times p(\gamma, \varphi')}{p(X, \varphi')} = \frac{p(X|\gamma, \varphi') \times p(\gamma, \varphi')}{p(X|\varphi') \times p(\varphi')} =$$

$$\frac{p(X|\gamma, \varphi')}{p(X|\varphi')} \times \frac{p(\gamma, \varphi')}{p(\varphi')} = \frac{p(X|\gamma, \varphi') \times p(\gamma|\varphi')}{\underbrace{p(X|\varphi')}_{\text{observed evidence}}} \propto \frac{p(X|\gamma, \varphi')}{\text{likelihood}} \times \frac{p(\gamma|\varphi')}{\text{prior}}.$$

- In Bayesian terms, the posterior probability distribution of the unknown parameter  $\gamma$  is proportional to the product of the likelihood and the prior.
- In probability terms, the posterior = likelihood × prior, divided by the observed evidence, in this case, a single spacetime data point,  $x_{i_o}$ .



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## Bayesian Inference Representation

- Spacetime analytics based on a single spacetime observation  $x_{i_o}$  can be thought of as a type of Bayesian prior-predictive or posterior-predictive distribution estimation problem
- Prior predictive distribution of a new data point  $x_{j_o}$ , marginalized over the *prior* – i.e., the sampling distribution  $p(x_{j_o}|\gamma)$  weight-averaged by the pure *prior* distribution:

$$p(x_{j_o}|\varphi') = \int p(x_{j_o}|\gamma) \times \frac{p(\gamma|\varphi')}{\text{prior distribution}} d\gamma.$$

- Posterior predictive distribution of a new data point  $x_{j_o}$ , marginalized over the *posterior*; i.e., the sampling distribution  $p(x_{j_o}|\gamma)$  weight-averaged by the *posterior* distribution:

$$p(x_{j_o}|x_{i_o}, \varphi') = \int p(x_{j_o}|\gamma) \times \frac{p(\gamma|x_{i_o}, \varphi')}{\text{posterior distribution}} d\gamma.$$

- The difference between these two predictive distributions is that
- The posterior predictive distribution is updated by the observation  $X = \{x_{i_o}\}$  and the hyperparameter,  $\varphi$  (phase aggregator),
  - The prior predictive distribution only relies on the values of the hyperparameters that appear in the prior distribution



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## Bayesian Inference Simulation

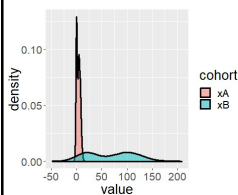
- Simulation example using 2 random samples drawn from mixture distributions each of  $n_A = n_B = 10K$  observations
  - 1)  $\{X_{A,i}\}_{i=1}^{n_A}$ , where  $X_{A,i} = 0.3U_i + 0.7V_i$ ,  $U_i \sim N(0,1)$  and  $V_i \sim N(5,3)$ , and
  - 2)  $\{X_{B,i}\}_{i=1}^{n_B}$ , where  $X_{B,i} = 0.4P_i + 0.6Q_i$ ,  $P_i \sim N(20,20)$  and  $Q_i \sim N(100,30)$ .
- The intensities of cohorts  $A$  and  $B$  are independent and follow different mixture distributions. We'll split the first cohort ( $A$ ) into training ( $C$ ) and testing ( $D$ ) subgroups, and then
  - 1) Transform all four cohorts into Fourier k-space,
  - 2) Iteratively randomly sample single observations from the (training) cohort  $C$ ,
  - 3) Reconstruct the data into spacetime using a single kime-magnitude value and alternative kime-phase estimates derived from cohorts  $B$ ,  $C$ , and  $D$ , and
  - 4) Compute the classical spacetime-derived population characteristics of cohort  $A$  and compare them to their spacekime counterparts obtained using a single  $C$  kime-magnitude paired with  $B$ ,  $C$ , or  $D$  kime-phases.



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## Bayesian Inference Simulation

Summary statistics for the original process (cohort  $A$ ) and the corresponding values of their counterparts computed using the spacekime reconstructed signals based on kime-phases of cohorts  $B$ ,  $C$ , and  $D$ . The estimates for the latter three cohorts correspond to reconstructions using a single spacetime observation (i.e., single kime-magnitude) and alternative kime-phase priors (in this case, kime-phases derived from cohorts  $B$ ,  $C$ , and  $D$ ).



	Spacetime	Spacekime Reconstructions (single kime-magnitude)		
Summaries	(A) Original	(B) Phase=Diff. Process	(C, training) Phase=True	(D, testing) Phase=Independent
Min	-2.38798	-3.798440	-2.98116	-2.69808
1 <sup>st</sup> Quartile	-0.89359	-0.636799	-0.76765	-0.76453
Median	0.03311	0.009279	-0.05982	-0.08329
Mean	<b>0.00000</b>	<b>0.000000</b>	<b>0.000000</b>	<b>0.000000</b>
3 <sup>rd</sup> Quartile	0.75772	0.645119	0.72795	0.69889
Max	3.61346	3.986702	3.64800	3.22987
Skewness	0.348269	0.001021943	0.2372526	0.31398
Kurtosis	<b>-0.68176</b>	<b>0.2149918</b>	<b>-0.4452207</b>	<b>-0.3270084</b>

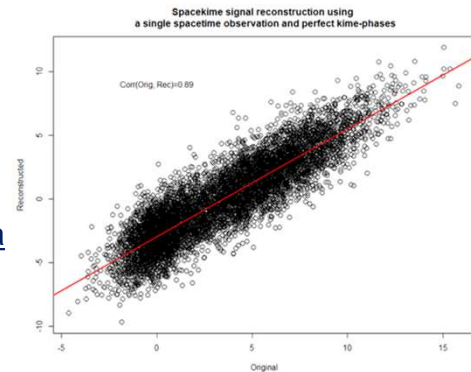


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## Bayesian Inference Simulation

The correlation between the original data ( $A$ ) and its reconstruction using a single kime magnitude and the correct kime-phases ( $C$ ) is  $\rho(A, C) = 0.89$ .

This strong correlation suggests that a substantial part of the  $A$  process energy can be recovered using only a single observation. In this case, to reconstruct the signal back into spacetime and compute the corresponding correlation, we used a single kime-magnitude (sample-size=1) and process  $C$  kime-phases.



**M**

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## Bayesian Inference Simulation

Let's demonstrate the Bayesian inference corresponding to this spacekime data analytic problem using a simulated bimodal experiment:

$$X_A = 0.3U + 0.7V, \text{ where } U \sim N(0,1) \text{ and } V \sim N(5,3)$$

Specifically, we will illustrate the Bayesian inference using repeated single spacetime observations from cohort  $A$ ,  $X = \{x_{i_o}\}$ , and varying kime-phase priors ( $\varphi = \text{phase aggregator}$ ) obtained from cohorts  $B$ ,  $C$ , or  $D$ , using different posterior predictive distributions.

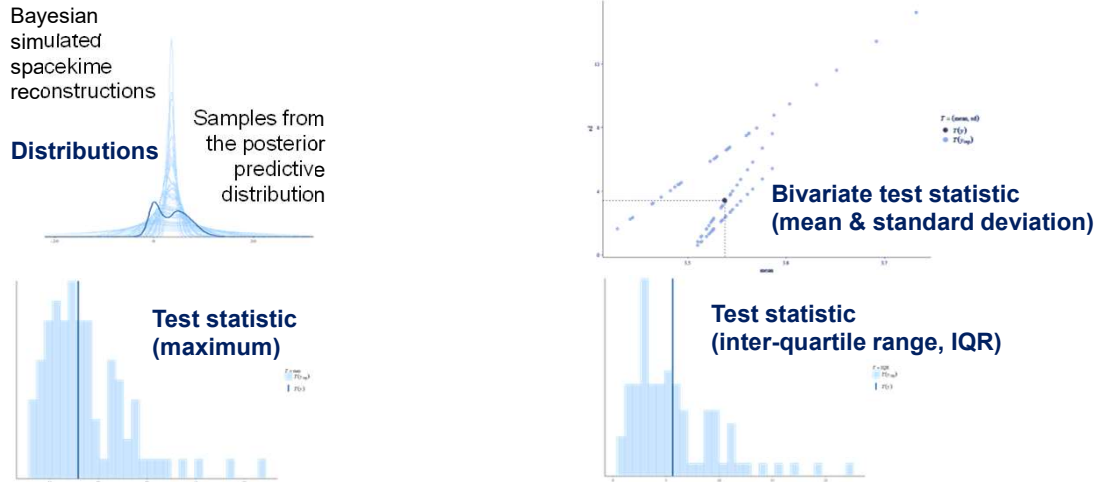
Relations between the empirical data distribution (**dark blue**) and samples from the posterior predictive distribution, representing Bayesian simulated spacekime reconstructions (**light-blue**). The derived Bayesian estimates do not perfectly match the empirical distribution of the simulated data, yet there is clearly information encoding that is captured by the spacekime data reconstructions.

This signal compression can be exploited by subsequent model-based or model-free data analytic strategies for retrospective prediction, prospective forecasting, ML classification, AI derived clustering, and other spacekime inference methods.

**M**

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# Bayesian Inference Simulation



Empirical data distribution (dark blue) & samples from the posterior predictive distribution Bayesian spacekime reconstructions (light-blue).



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# Part 2: Hands-on Spacekime Analytics Tutorial

[TCIU/Spacekime Analytics Tutorial: Basic TCIU Protocol for Predictive Spacekime Analytics using Longitudinal Data](https://www.socr.umich.edu/TCIU/HTMLs/Chapter6_TCIU_Basic_SpacekimePredictiveAnalytics.html)

1 Preliminary setup

2 Longitudinal Data Import

3 Time-series graphs

3.1 Interactive time-series visualization

3.1.1 Example fMRI(x=4, y=42, z=33, t)

4 kime-series/kime-surfaces (spacekime analytics protocol)

4.1 Pseudo-code

4.2 Function main step: Time-series to Kime-surfaces Mapping

4.2.1 Generate the kime-phases

4.2.2 Structural Data Preprocessing

4.2.3 Intensity Data Preprocessing

4.2.4 Generate `ijk` labels

4.2.5 Cartesian space interpolation

4.2.6 Cartesian representation

4.2.7 Generate a long data-frame

4.2.8 Display kime-surfaces

4.3 Analytical Time-series to Kime-Surface Transformation (kernel)

[SOCR](#)   [TCIU Website](#)   [TCIU Github](#)

## Spacekime Analytics (Time Complexity and Inferential Uncertainty) Code

[Show All Code](#)  
[Hide All Code](#)

### Basic TCIU Protocol for Predictive Spacekime Analytics using Longitudinal Data

SOCR Team  
05/13/2024

This [Spacekime TCIU Learning Module](#) presents the core elements of spacekime analytics including:

- Import of repeated measurement longitudinal data,
- Numeric (stitching) and analytic (Laplace) kime-surface reconstruction from time-series data,
- *Forward prediction modeling* extrapolating the process behavior beyond the observed time-span  $[D, T]$ ,
- *Group comparison discrimination* between cohorts based on the structure and properties of their corresponding kime-surfaces. For instance, statistically quantify the differences between two or more groups,
- *Unsupervised clustering and classification* of individuals, traits, and other latent characteristics of cases included in the study,
- Construct low-dimensional visual representations of large repeated measurement data across multiple individuals as pooled kime-surfaces (parameterized 2D manifolds),
- Statistical comparison and quantitative contrasting of kime-surface differences.

### 1 Preliminary setup

TCIU and other `R` package dependencies ... Show

### 2 Longitudinal Data Import

In this case, we are just loading some exemplary fMRI data, which is [available here](#). Show

[https://www.socr.umich.edu/TCIU/HTMLs/Chapter6\\_TCIU\\_Basic\\_SpacekimePredictiveAnalytics.html](https://www.socr.umich.edu/TCIU/HTMLs/Chapter6_TCIU_Basic_SpacekimePredictiveAnalytics.html)



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